Appendix A

Glossary

Absolute magnitude Apparent magnitude a star would have at a distance of 10

parsecs (32.6 LY).

Absorption lines Dark lines superimposed on a continuous electromagnetic

spectrum due to absorption of certain frequencies by the medium

through which the radiation has passed.

AM Amplitude modulation. Imposing a signal on transmitted energy

by varying the intensity of the wave.

Ångstrom Unit of length equal to 10^{-10} m. Sometimes used to measure

wavelengths of visible light. Now largely superseded by the

nanometer (10⁻⁹ m).

Aphelion The point in a body's orbit (Earth's, for example) around the sun

where the orbiting body is farthest from the sun.

Apogee The point in a body's orbit (the moon's, for example) around

Earth where the orbiting body is farthest from Earth.

Apparent magnitude Measure of the observed brightness received from a source.

Astrometry Technique of precisely measuring any wobble in a star's

position that might be caused by planets orbiting the star and

thus exerting a slight gravitational tug on it.

Astronomical horizon The hypothetical interface between Earth and sky, as would be

seen by the observer if the surrounding terrain were perfectly

flat.

Astronomical unit Mean Earth-to-sun distance, approximately 150,000,000 km.

Atmospheric window Property of Earth's atmosphere that allows only certain wave-

lengths of electromagnetic energy to pass through, absorbing all

other wavelengths.

AU Abbreviation for astronomical unit.

Azimuth In the horizon coordinate system, the horizontal angle from some

arbitrary reference direction (north, usually) clockwise to the

object circle.

Background radiation Electromagnetic radiation originating from no specific location.

Beam width The angle within which an antenna receives radio waves.

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Binary stars Two stars that are so close as to orbit around a common center of

gravity.

Black dwarf One possible final stage in the evolution of a star, in which all

the energy is exhausted and it no longer emits radiation.

Black hole A region of space surrounding a very massive collapsed star.

Gravity is so intense in a black hole that the escape velocity is equal to or greater than the speed of light, thus no radiation can

escape.

Blackbody A hypothetical object that is a perfect radiator, absorbing and re-

emitting all radiation that impinges upon it.

Blazar Short for BL Lacertae object: a class of galaxies whose light is

polarized and has a non-thermal spectrum, and whose brightness

fluctuates dramatically.

Blue shift Apparent shortening of the wavelength, due to Doppler effect, of

radiation received from a source in motion toward the observer.

Bowshock The upstream (closest to the sun) boundary of a planet's mag-

netosphere, interfacing with the solar wind.

Brightness spectrum Variation of radiation brightness with frequency.

Brightness Also referred to as surface brightness. Power of radiation

received per unit area per unit solid angle per unit frequency interval. Note that this is not the common usage of "brightness." What is commonly referred to as brightness is a characteristic of the source of the radiation (rather than the radiation received)

and is what astronomers call luminosity.

Celestial equator A great circle on the celestial sphere, concentric with Earth's

equator.

Celestial latitude In the ecliptic coordinate system, the angle of an object in the

sky north or south (±90°) of the ecliptic plane.

Celestial longitude In the ecliptic coordinate system, the angle of an object in the

sky eastward along the ecliptic from the vernal equinox (0-360°).

Celestial poles Points about which the celestial sphere appears to rotate; inter-

sections of the celestial sphere with Earth's axis.

Celestial sphere Apparent sphere of the sky, whose diameter is the distance from

Earth to the farthest observable object.

Cepheid A type of regular variable star that has a short period of a few

days to a few weeks. It has been found that in these stars, the

longer the period, the more luminous the star.

Chromosphere The part of the sun's atmosphere lying above the photosphere

and below the outer corona.

Corona The outer atmosphere of the sun. Visible to the naked eye only

during a total solar eclipse.

Cyclotron radiation Electromagnetic radiation emitted when charged particles are

moved within a magnetic field at non-relativistic speeds (that is,

not close to the speed of light).

Declination Angular distance north or south of the equator of some object in

the sky, measuring along an hour circle passing through that

object.

Discrete source A sources of radiation whose direction can be identified. Dis-

crete sources may be further classified as point sources, localized

sources, and extended sources.

Doppler effect Apparent change in wavelength of the radiation from a source

due to its motion relative to the observer.

Ecliptic coordinate system Coordinate system using the plane of the ecliptic as the refer-

ence.

Electromagnetic spectrum The full range of wavelengths or frequencies of electromagnetic

radiation.

Electromagnetic radiation Radiation consisting of waves propagated through the building

up and breaking down of electric and magnetic fields; include radio, infrared, light, ultraviolet, x-rays, and gamma rays.

Elevation In the horizon coordinate system, the angle upward from the

horizon to an object in the sky.

Emission lines Discrete, bright lines in the spectrum.

Epoch A date chosen for reference purposes in quoting astronomical

coordinates, to account for the slight changes in the apparent positions of objects in the sky due to Earth's precession.

Equator A great circle on Earth 90° from its poles.

Equatorial coordinate system A coordinate system using Earth's equator as the plane of

reference.

Extended source Discrete emitter of radiation that covers a relatively large part of

the sky, for example, the Milky Way galaxy. The terms localized and extended are relative and depend on the angular resolu-

tion of the telescope observing them.

Faraday rotation Rotation of an electromagnetic wave's polarization as it passes

through a magnetic field parallel to the propagation of the wave

in a medium.

Field The effect of forces, such as gravity and magnetism, that act on

distant objects.

Flare star Faint red dwarf stars that may brighten up by several magnitudes

over a few minutes, and fade back to their usual brightness

within an hour.

FM Frequency modulation. Imposing a signal on transmitted energy

by varying the frequency of the wave.

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Galactic coordinate system A coordinate system using the plane of the Milky Way galaxy

as the reference.

Gravitational lens An effect produced by the bending of radiation as it passes

within the gravitational field of another object. May appear to an observer on Earth as one or more duplicate images of the

same source.

Horizon coordinate system A coordinate system using a plane through the observing point

parallel to the horizon as the reference.

Horizon mask A diagram that maps in silhouette the horizon in 360° as

actually seen by the observer.

Hour angle The elapsed time since an object in the sky crossed the merid-

ian point of the observer.

Hour circle (object's)

A great circle on the celestial sphere that passes through both

poles and the object being observed in the sky.

Index of refraction Ratio of the speed of electromagnetic radiation in a vacuum to

the speed of electromagnetic radiation in a given medium.

Intensity Power of electromagnetic radiation received per unit area.

Interferometry In radio astronomy, use of more than one radio telescope to

enhance the resolution of the radio image from a source.

Inverse-square law The amount of electromagnetic energy reaching a given point

in a given unit of time decreases in proportion to the square of

the distance from the source of the energy.

ISM Interstellar medium.

Kelvin A unit of absolute temperature. In the Kelvin temperature

scale, 0 is absolute zero, the temperature where molecular motion ceases and the material has no kinetic energy. Water

freezes at 273 K.

Light year Distance light (electromagnetic energy) travels in a vacuum

during one year, or $9.46 \times 10^{12} \text{ km}$.

Localized source A discrete radiation source of very small extent. A single star

is normally considered a localized source.

LY Abbreviation for light year.

Magnetosphere Region around a planet in which its magnetic field dominates

the interplanetary field carried by the solar wind.

Maser (astronomical) For Microwave-amplified Stimulated Emission of Radiation.

An ultracompact site in molecular clouds (usually of water vapor, hydroxl radicals, silicon monoxide, or methanol) in the interstellar medium or very near long-period variable stars, in which certain molecular emissions lines are enormously

amplified.

Meridian A great circle on the celestial sphere that passes through both

poles and the zenith.

Nadir The point on the celestial sphere immediately below the ob-

server. Directly opposite the overhead point, or zenith.

Nanometer One thousand-millionth of a meter, 10^{-9} m. Commonly used to

measure wavelengths in the visible light and UV ranges.

Nebula Cloud of interstellar gas or dust.

Neutron star A star of extremely high density but low luminosity, composed

almost entirely of neutrons.

Non-thermal emissions Electromagnetic radiation produced by synchrotron radiation,

maser line emissions from atoms and molecules, or other mecha-

nisms not related to temperature.

Nutation (Earth's) A slight, slow nodding of Earth's axis (about 9 arcsecs to either

side of its mean position in about 18 years 220 days). This motion is apart from precession of the axis and is caused by the gravitational pull of the moon on Earth's equatorial bulge as the moon moves slightly above and slightly below the ecliptic in its

orbit.

Object circle In the horizon coordinate system, the vertical circle through a

celestial object and the zenith.

Oblateness The flattening of a sphere. Earth is oblate such that its diameter

from pole to pole is 43 km less than its equatorial diameter.

Occultation Passing of one celestial body in front of another as observed

from Earth.

Optical astronomy Study of the extraterrestrial universe using primarily visible light

observations.

Optical window The characteristic of Earth's atmosphere that allows visible light

to pass through.

Parsec The distance of an object at which 1 AU would subtend one

arcsec; 1 parsec = 3.26 light years.

Perigee The point in a body's orbit (the moon's, for example) around

Earth when the orbiting body is closest to Earth.

Perihelion The point in a body's orbit (Earth's, for example) around the sun

when the orbiting body is closest to the sun.

Phase Angular distance between peaks or troughs of two wave forms of

similar frequency.

Photon astronomy Study of the extraterrestrial universe using observations of the

entire electromagnetic spectrum.

Photosphere The region of the sun (or any star) from which continuous

spectrum radiation escapes into space; the visible disk of the sun.

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Plasma A hot, ionized gas.

Point source An idealized discrete source of radiation that subtends an infini-

tesimally small angle.

Polarization The direction of the electric vector of an electromagnetic wave.

Polarization may be linear, circular, random, elliptical, or any

combination.

Precession (Earth's) A slow, conical motion of Earth's axis of rotation, similar to the

wobble of a spinning top.

Prime vertical In the horizon coordinate system, the great circle through the

east and west points and the zenith.

Pulsar A neutron star radio source that emits in rapid, regular pulses.

Quasar A very distant, very luminous source of visible and radio energy.

May be the oldest and most distant objects in the universe.

Radio astronomy Study of the extraterrestrial universe using observations of radio

frequency radiation, rather than visible light.

Radio galaxy Galaxy that is a strong emitter of radio energy, perhaps emitting

10⁵ - 10⁶ times more radio energy than a "normal" galaxy.

Radio star A star that is a strong emitter of radio frequency radiation.

Radio window The property of Earth's atmosphere that allows certain wave-

lengths of electromagnetic radiation in the radio range to pass

through.

Recombination lines Emission and absorption lines seen when oppositely charged

ions recombine to an electrically neutral state, producing a highly excited neutral atom, with electrons transitioning between

states, emitting and absorbing photons.

Red shift Apparent lengthening of the wavelength of radiation received

from a source. May be caused by Doppler effect, the expansion

of space, or gravitational effects of very massive objects.

Right ascension In the equatorial coordinate system, the angle (usually in hours,

minutes, seconds) measured eastward along the celestial equator from the vernal equinox to the hour circle passing through an

object in the sky.

R₁ Abbreviation for one Jupiter radius (about 5 million km). Used

to describe distances in the Jupiter system.

RR Lyrae variable Regular variable stars whose periods are very short (between

about 1.25 and 30 hours).

S-band Range of radio frequencies of about 2-4 GHz, or wavelengths of

15 - 7.5 cm.

Scintillation Effect produced by phase shifting of electromagnetic waves from

a discrete source as they pass through Earth's atmosphere. In the

visible light range, perceived as twinkling of stars.

SETI Search for extraterrestrial intelligence. Using radio telescopes,

astronomers "listen" for artificially modulated radio frequency signals from parts of the universe most "likely" to spawn life as

we know it.

Sidereal day

The time required for Earth to revolve 360° with respect to a

celestial object outside the solar system. About 23 hours 56

minutes duration in terms of solar time.

Solar day

The time required for Earth to revolve 360° with respect to the

sun.

Solar flare Brilliant outbreak in the sun's outer atmosphere, usually associ-

ated with active groups of sunspots.

Spectral power Power of electromagnetic radiation per unit of frequency band-

width.

Stellar wind The outflow of gas from a star, sometimes at speeds as high as

hundreds of kilometers per second.

Sunspot A temporary cool region in the solar photosphere that appears

dark by contrast against the surrounding hotter photosphere.

Superluminal velocity Apparent motion of an object at greater than light speed;

apparency is caused by a projection effect due to the motion's

vector component toward Earth.

Supernova A star that experiences a cataclysmic explosion, sending much of

its material into space.

Surface brightness See *brightness*.

Synchrotron radiation Radiation emitted by charged particles being accelerated in

magnetic fields and moving at near the speed of light.

Thermal emissions Radiation emitted due to an object's temperature (for example,

blackbody radiation) or by an ionized gas.

Vernal equinox The point on the celestial sphere where the sun crosses the

celestial equator from south to north. The date of this crossing

(about March 21) is also called the vernal equinox.

X-band Range of radio frequencies of about 8-12 GHz, or wavelengths

of 3.75 - 2.4 cm.

Zenith The point on the celestial sphere directly overhead the observer.

Appendix B

References and Further Reading

Books:

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^{*} Recommended reading (i.e., technical level approximately the same as this workbook).

World-Wide Web Pages:

National Radio Astronomy Observatory Home Page, http://www.nrao.edu.

Water Masers in the Direction of the Galactic Center,

http://www.cv.nrao.edu/~eschulma/h2o.html.

What is the VLA?

http://www.nrao.edu/doc/vla/html.VLAintro.shtml

JPL Molecular Spectroscopy Home Page,

http://spec.jpl.nasa.gov

JPL Deep space Network Home Page,

http://deepspace.jpl.nasa.gov/dsn

Basics of Space Flight Learner's Workbook,

http://www.jpl.nasa.gov/basics

SETI (Search for Extraterrestrial Intelligence) Institute Home Page,

http://www.seti-inst.edu

Multiwavelength Atlas of Galaxies,

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http://www-wilson.ucsd.edu/education/spectroscopy/spectroscopy.html

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Video:

Powers of Ten. Copy available at AVSTC.

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| Chapter | Page | Subject | Source |
|---------|-------|-----------------------------------|--------------------------------|
| 1 | 4 | Jansky's antenna | Morrison et al., 162 |
| 1 | 5 | Reber's radio telescope | Morrison et al., 162 |
| 2 | 10 | Wavelength-frequency relationship | Morrison et al., 108 |
| 2 | 11 | Inverse-square law | Morrison et al., 109 |
| 2 | 13 | EM spectrum, general | Morrison et al., 110 |
| 2 | 14 | EM wavelenth/freq. chart | Doody & Stephan, 45 |
| 2 | 16 | Linear polarization | Halliday & Resnick, 1147 |
| 2 | 16 | Wave cross-section | Halliday & Resnick, 981 |
| 2 | 17 | Circular polarization | Halliday & Resnick, 1164 |
| 3 | 22 | Brightness spectrum curves | Kraus, 322 |
| 3 | 24 | Hydrogen 21-cm line | Morrison et al., 422 |
| 3 | 26 | Synchrotron radiation | Morrison et al., 579 |
| 3 | 27 | Thermal vs. non-thermal | Heiserman, 94 |
| 4 | 30 | Atmospheric windows | Composite "original" |
| 4 | 31 | Kirchhoff's Laws | Morrison et al., 119 |
| 4 | 32 | Hydrogen atom | Morrison et al., 116 |
| 4 | 34 | Reflection | Doody & Stephan, 52 |
| 4 | 35 | Antenna types | Doody & Stephan, 53 |
| 4 | 36 | Air-glass-air refraction | Doody & Stephan, 54 |
| 4 | 36 | Atmospheric refraction | Doody & Stephan, 55 |
| 4 | 36 | Phase | Doody & Stephan, 55 |
| 4 | 38 | Faraday rotation | Chen, 133 |
| 5 | 39 | Doppler effect | Doody & Stephan, 51 |
| 5 | 41 | Gravitational lensing | Morrison et al., 582 |
| 5 | 42 | Superluminal velocity | Morrison et al. (6th Ed.), 581 |
| 6 | 46 | Sources classifications | Original |
| 6 | 47 | Four views of sky | Kaufmann, 91 |
| 6 | 49 | Pulsar | Kraus, 9-27 |
| 6 | 50 | Sun | Original |
| 6 | 51 | Optical vs. radio solar flares | Shields, 17 |
| 6 | 53 | Radio view of Milky Way | Moore, 188 |
| 6 | 55 | Jupiter magnetosphere | Morrison et al., 285 |
| 7 | 60 | Terrestrial coordinates | Doody & Stephan, 18 |
| 7 | 61 | Solar vs. sidereal time | Morrison et al., 83 |
| 7 | 62 | Precession of Earth axis | Morrison et al., 72 |
| 7 | 63 | Horizon coordinate system | Kraus, 2-22 |
| 7 | 64 | Horizon mask for GAVRT | NASA Goldstone DSSC |
| 7 | 65 | Equatorial coordinate system | Original |
| 7 | 66 | Precessional path | Doody & Stephan, 22 |
| 7 | 67 | Antenna mount types | Doody & Stephan, 22 |
| 7 | 68 | Galactic coordinate system | Kraus, 2-31 |
| 8 | 72-74 | Universe in 6 steps | Kraus, 2-13 thru 18 |